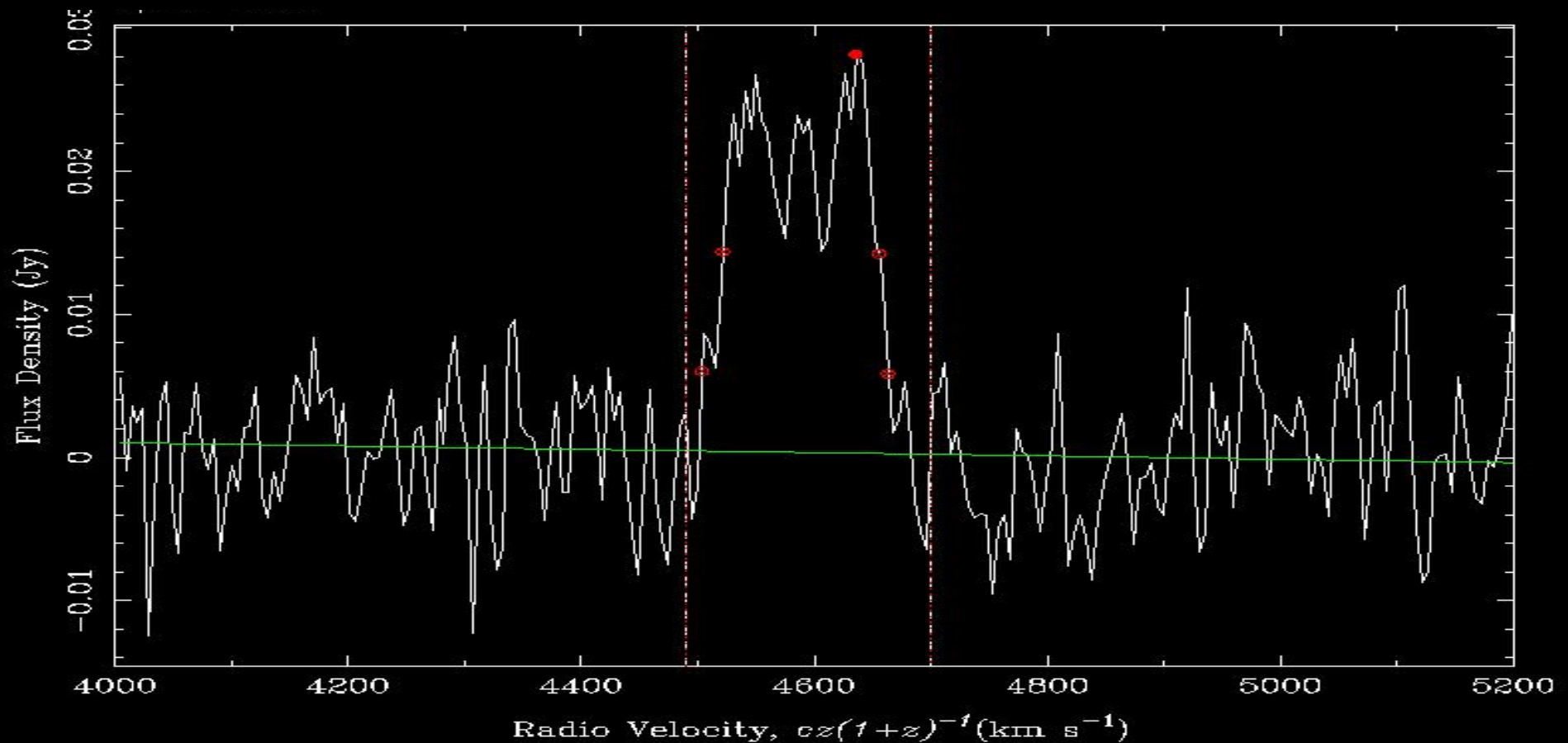


ALFA Zone of Avoidance (ZOA) Survey

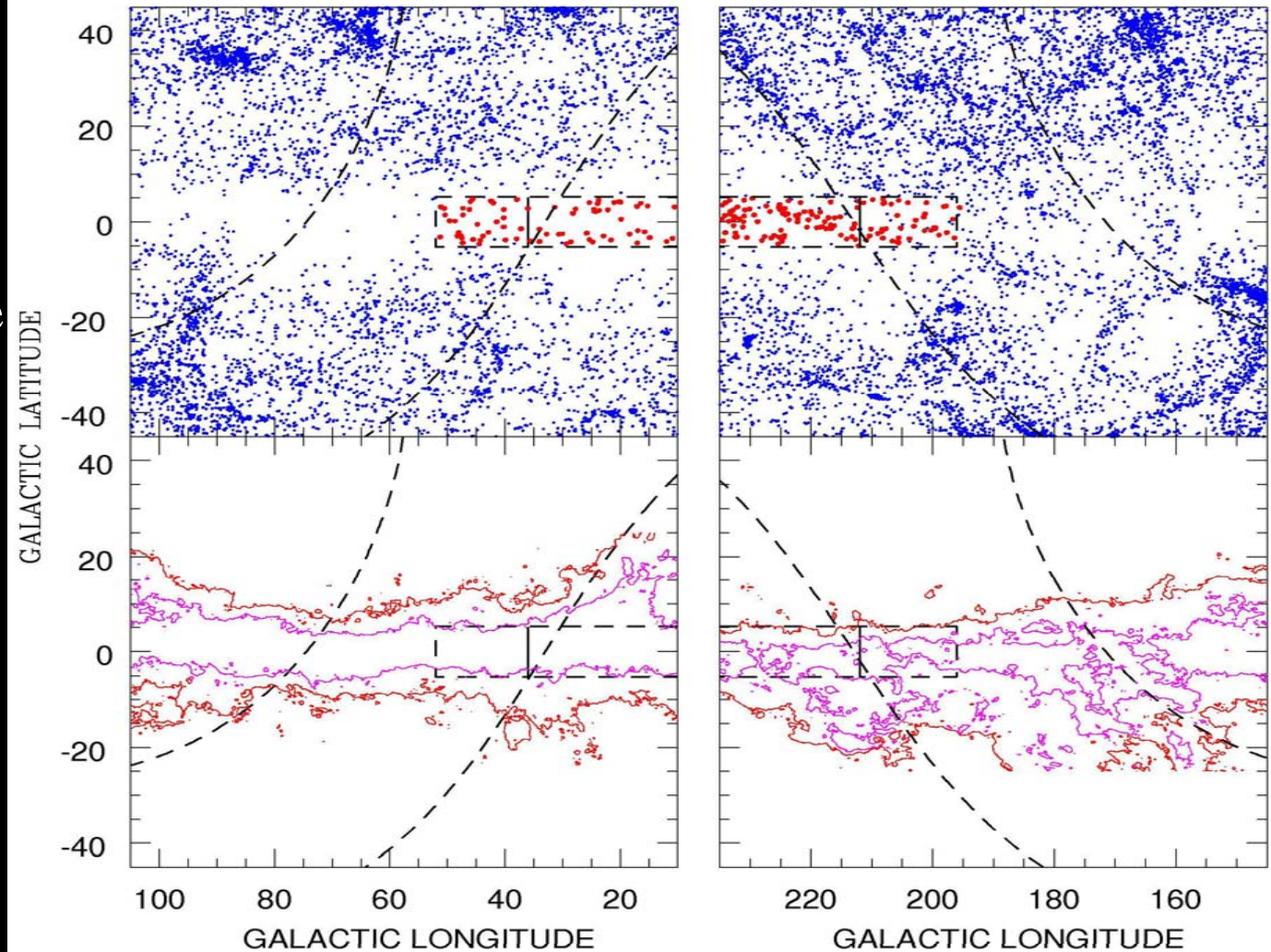
UNM	Trish Henning	ATNF	Baerbel Koribalski
NRL	Chris Springob		Erik Muller
NAIC	Barbara Catinella		Lister Staveley-Smith
	Emmanuel Momjian	U. PR	Carmen Pantoja
CfA	Karen Masters	U. MI	Mary Putman
	Jessica Rosenberg	U. MA	Stephen Schneider



- ALFA ZOA Survey
 - Commensal with GALFA and PALFA surveys of ZOA region
- Why survey the ZOA at HI?
 - Scientific motivations similar to that of ALFALFA
 - e.g., HIMF/missing satellite problem, study of HVCs, etc.
 - like ALFALFA, the sky being surveyed is unbiased; only determined by orientation of our galaxy
 - Mapping large scale structure in the ZOA
 - ZOA has been narrowed by deep optical/IR searches, but both fail in regions of high extinction and stellar confusion
 - HI can still be seen in these regions
 - 3d connectivity of LSS across ZOA
 - Structures like the Pisces-Perseus Supercluster and the Great Attractor are partially hidden by Milky Way

- HI Parks Zone of Avoidance Survey (HIZOA)--surveyed mostly in southern hemisphere to 6 mJy/beam rms

ZOA in the Arecibo sky. **TOP Blue:** galaxies in the literature within 12,000 km/s in and near regions accessible to AO. **Red:** HIZOA galaxies. **BOTTOM Red:** $A_B=1$ mag of extinction contour; **Blue:** $A_B=3$ mag of extinction contour.



- ZOA plan

- GALFA

- map within 10 degrees of plane, across 45 degrees longitude of inner Galaxy, 45 degrees longitude of outer Galaxy (~5 sec integration time)

- PALFA

- map within 5 degrees of plane, across 45 degrees longitude of inner Galaxy (268 sec integration time), 45 degrees longitude of outer Galaxy (134 sec integration time)

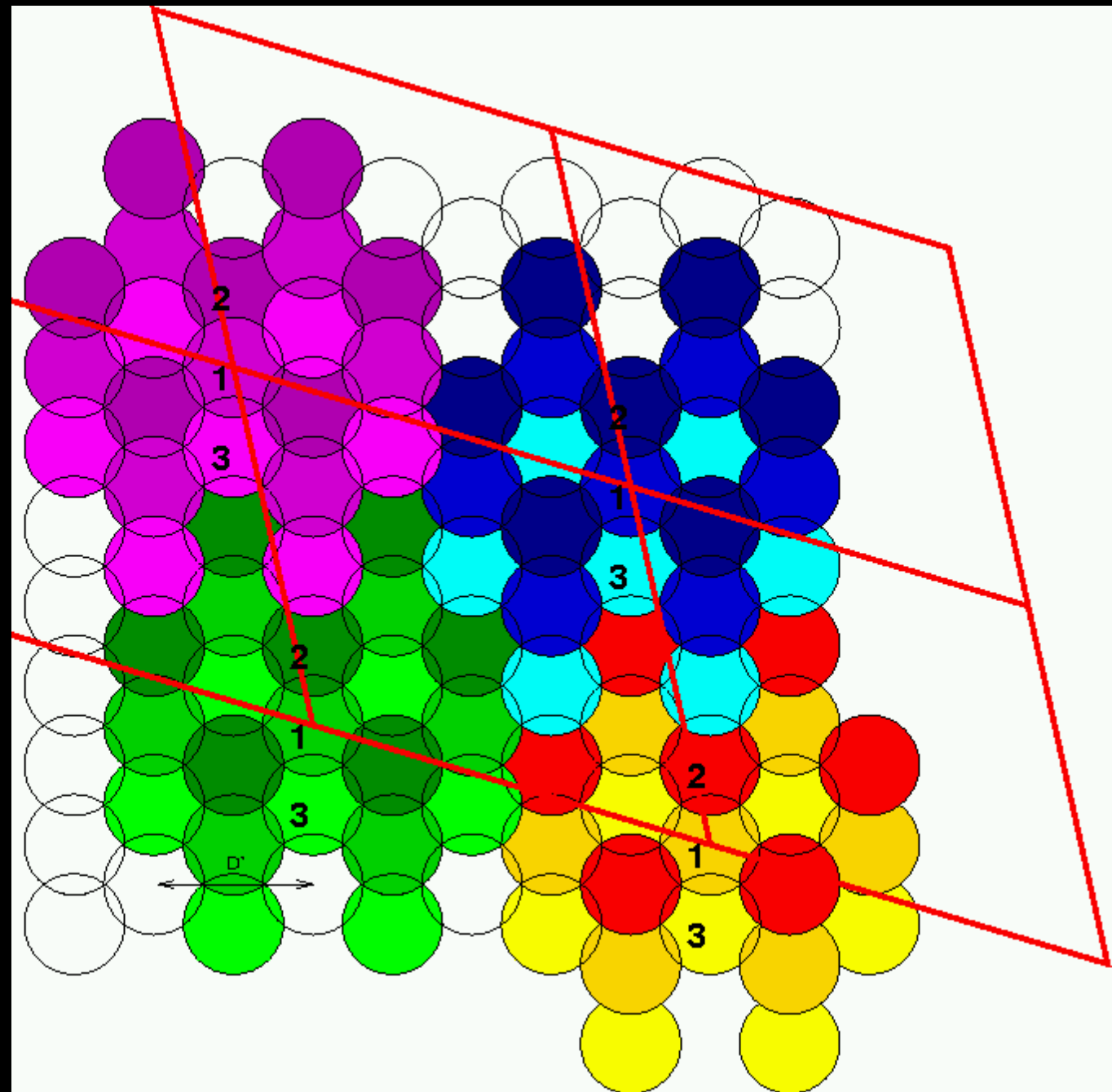
- PALFA-ZOA

- Leapfrog tiling

- point and stare at one spot in the sky, then move beam to next part of sky

- Strategy: **sparse (1/3) sampling out from $|b|=0$ to higher latitudes, then fill in rest of sky**

- PALFA survey has already started; proposal for 1st year of PALFA-ZOA was given an 'A' grade earlier this year, but we can't start until PALFA spectrometer or EALFA spectrometer is ready (early 2007)



• GALFA-ZOA

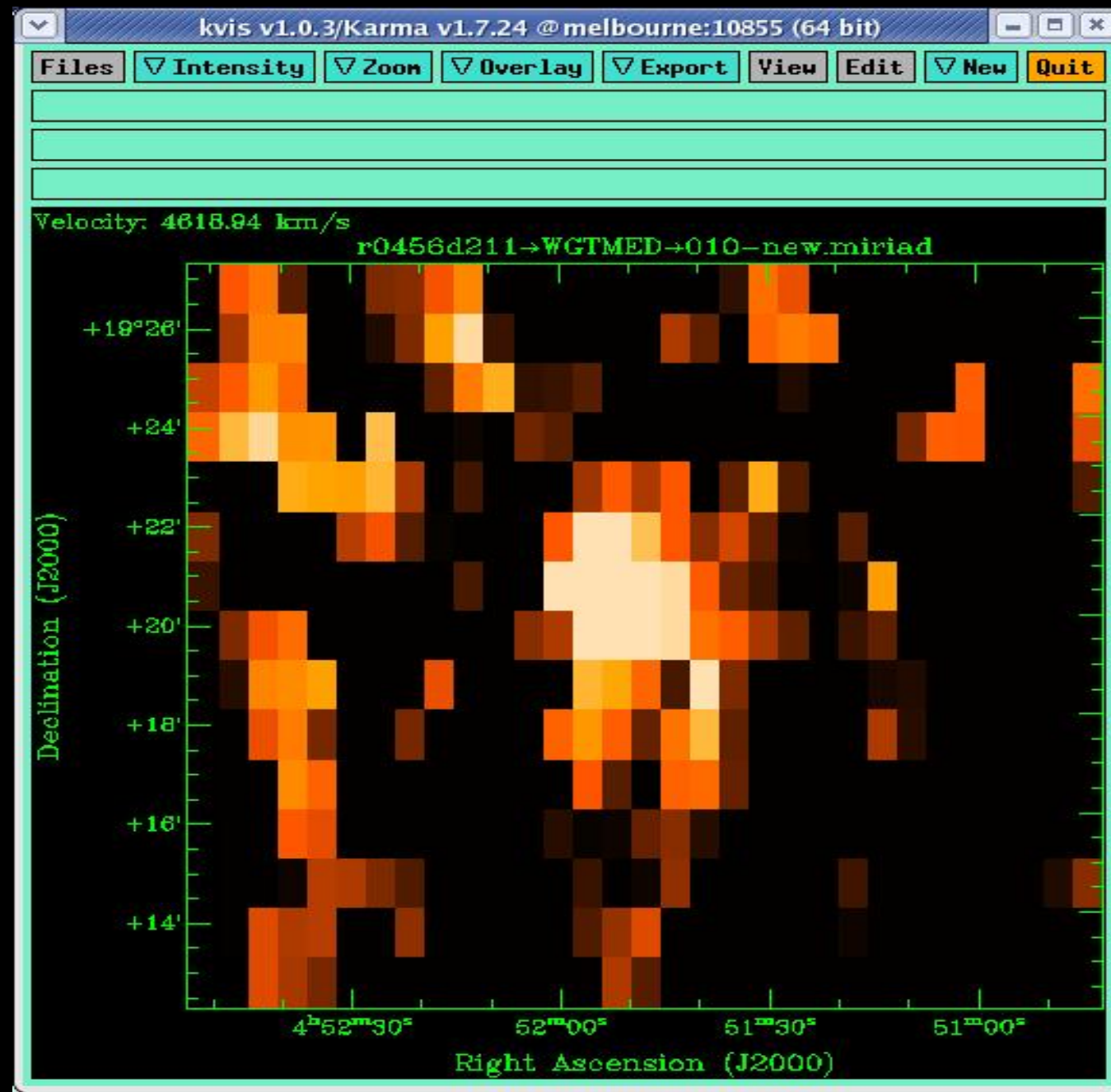
– Status

- We have covered $\sim 140 \text{ deg}^2$ in precursor observations in 2005-2006
 - data is in various stages of reduction
- Proposal for inner galaxy part of real survey was given an 'A' grade in January, but corresponding GALFA proposal awaits critical review

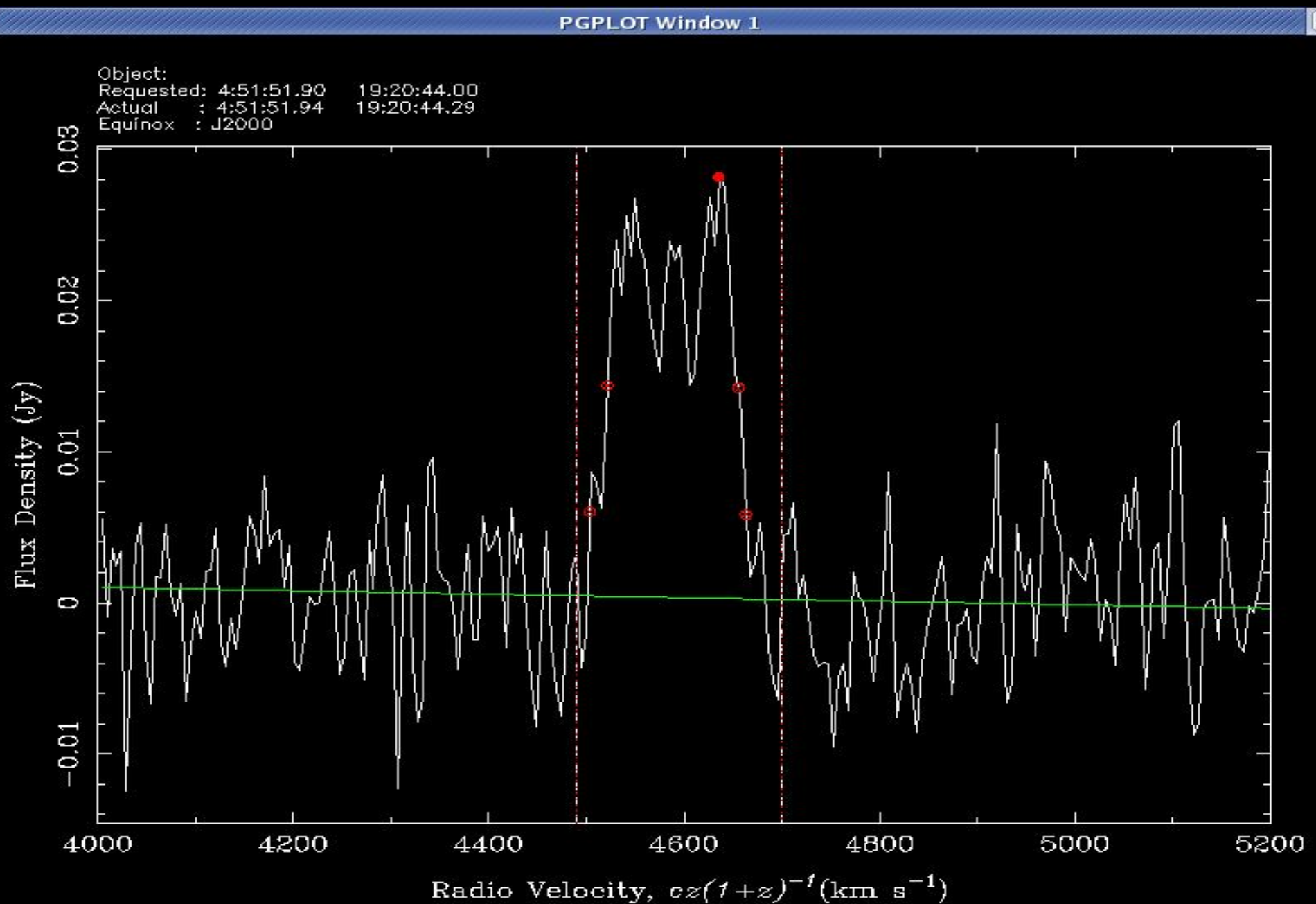
– Observational setup

- basketweave method: telescope slews back and forth in declination while RA drifts with sky; covers ~ 4 degrees in 3 minutes each direction
- we observe with WAPPS, while GALFA observes with their own spectrometer
- bandwidth: $-2600-18,400 \text{ km/s}$
- spectral resolution: $\sim 5 \text{ km/s}$ (HIZOA: $\sim 13 \text{ km/s}$)
- angular resolution: $\sim 3.5'$ (HIZOA: $\sim 15.5'$)

- **Data reduction pipeline: closely follows that of AGES**
 - Livedata and Gridzilla get you gain-corrected, calibrated, gridded dataset
 - Look through data cube for galaxies by eye, building up catalog of possible detections, and then going back to confirm later by inspecting individual profiles.



- Measure HI parameters using Miriad.
- Check coordinates against possible counterparts in NED.



The Near Future

- **GALFA-ZOA**
 - refinement of HI parameter extraction procedure
 - completion of precursor data reduction
 - beginning of real survey: date TBD???
- **PALFA-ZOA**
 - development of PALFA-ZOA pipeline
 - survey to begin early 2007