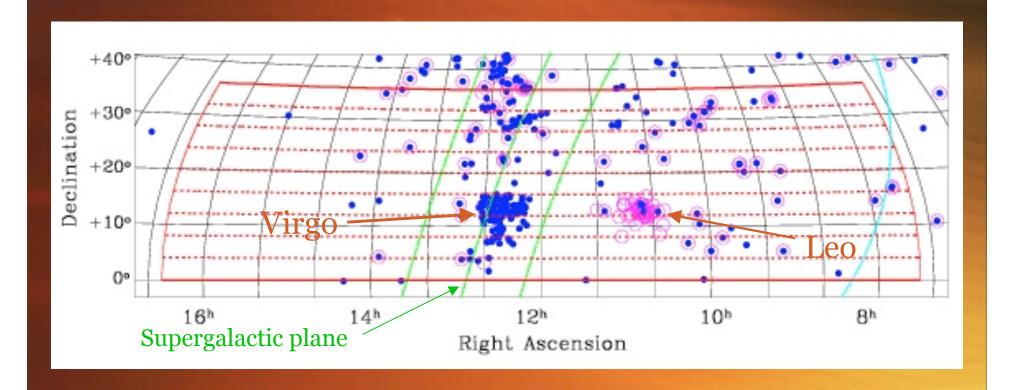
# ALFALFA Survey of the Leo Region

Sabrina Stierwalt
ALFALFA Workshop
June 23, 2006
Cornell University

# ALFALFA: Spring Sky



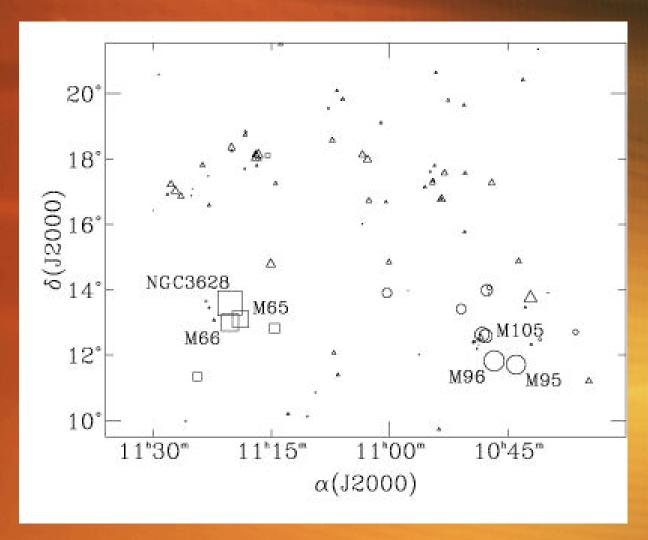
Leo Group

• RA: 09<sup>h</sup>45<sup>m</sup> to 11<sup>h</sup>45<sup>m</sup>

• Dec: 8° to 16°

## Quick Tour of the Leo Region

- Likely M66 group member
- C Likely M96 group member
- △ Likely Leo II member



Plot from Schneider's Leo Group in Encyclopedia of Astronomy & Astrophysics

## Leo I Group

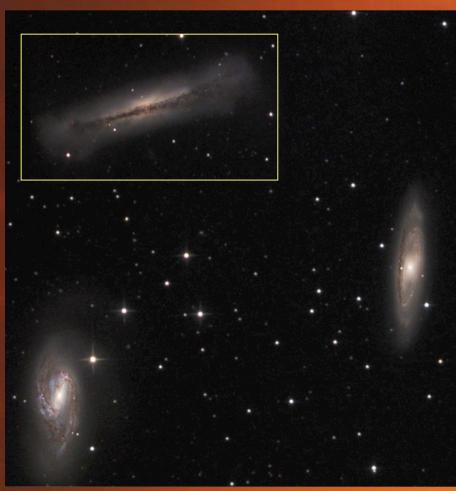
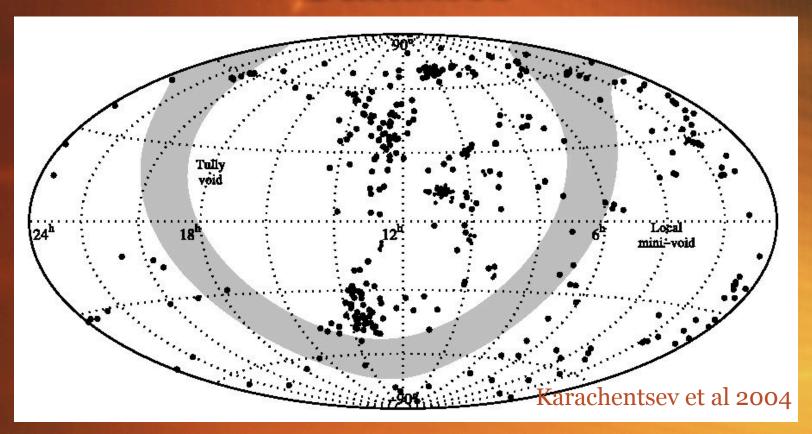


Image Credit: NOAO

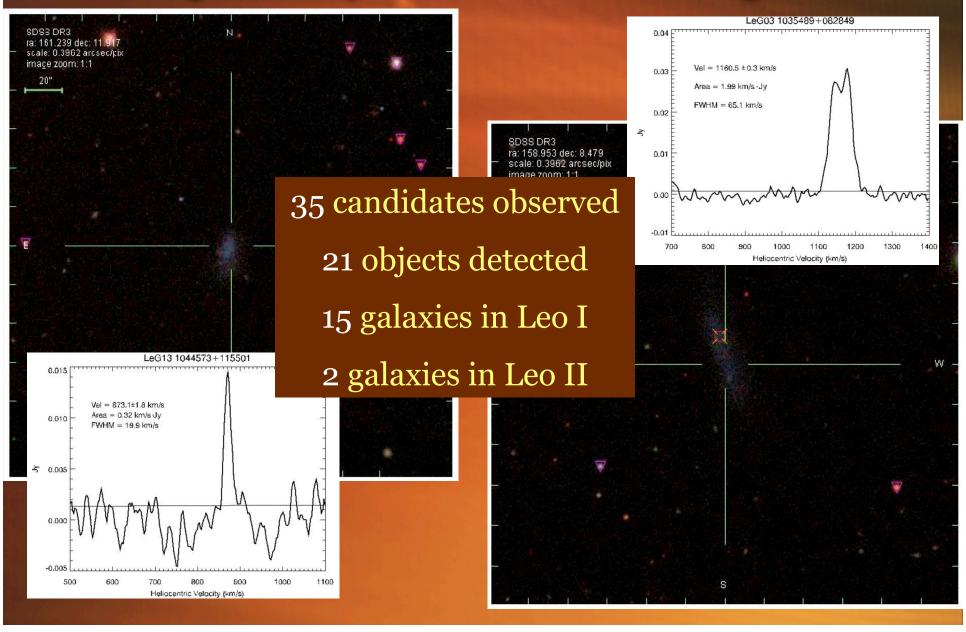
- •Proximity allows for study of low-mass, low-surface brightness galaxies
- •Attractive opportunity for studying the HIMF in an intermediate density environment
- •Interesting environment because has low velocity dispersion (~130 km/s) but a local density enhancement high enough to support presence of E/So galaxies

## K & K's Catalog of Nearby Galaxies



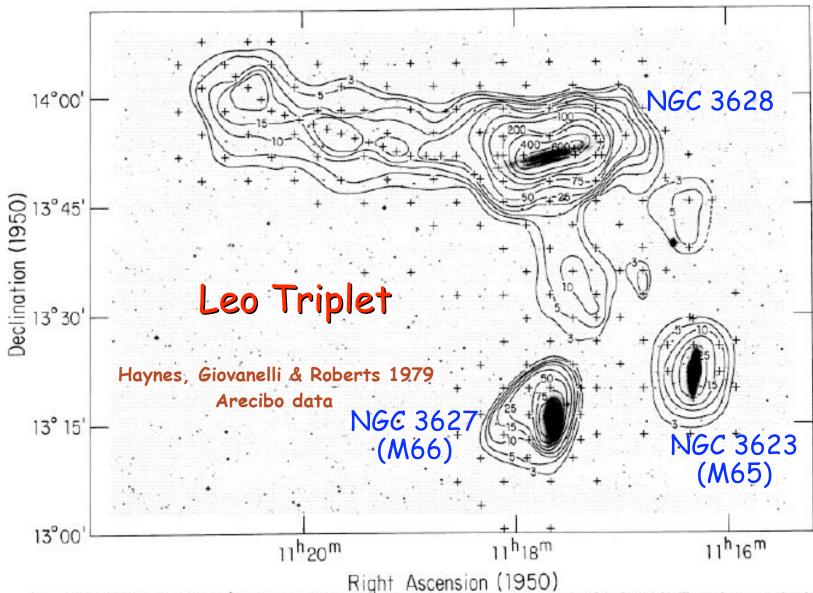
- 451 galaxies with D  $\leq$  10 Mpc or  $v_{LG} < 550$  km/s
- 35 dwarf galaxy candidates identified via visual inspection of POSS-II/ESO/Serc plates

## Optically-Selected Sample



### HI-Selected Sample (ALFALFA)

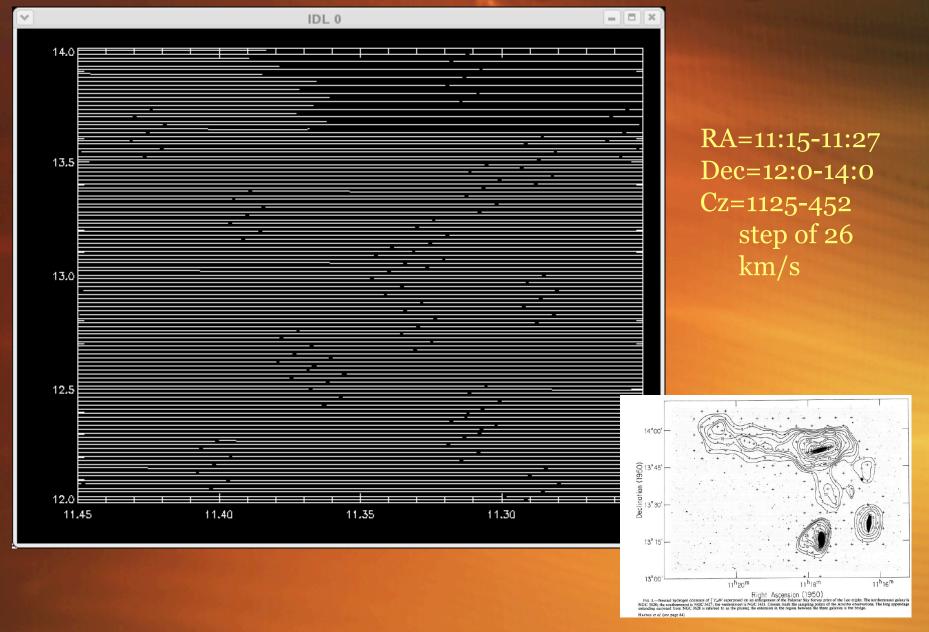
- 208 hours of observing time in 240 deg<sup>2</sup>
- Of the 21 (LBW-detected) dwarf galaxy candidates from the optically-selected sample, 15 found in ALFALFA (6 required longer integration times)
- List of many new dwarf galaxy candidates being added to every day

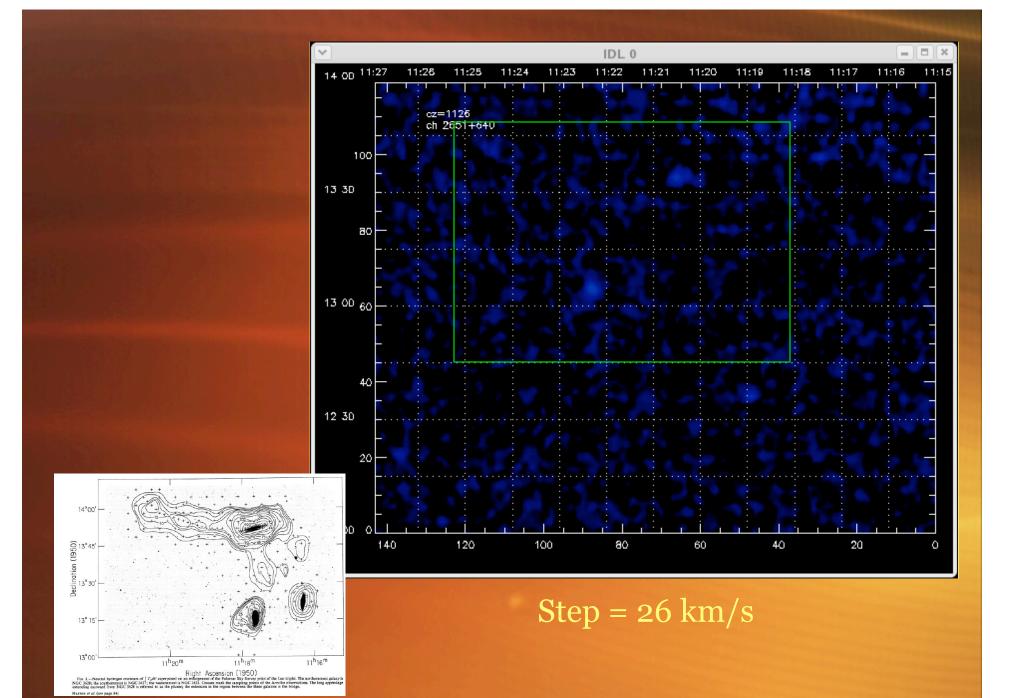


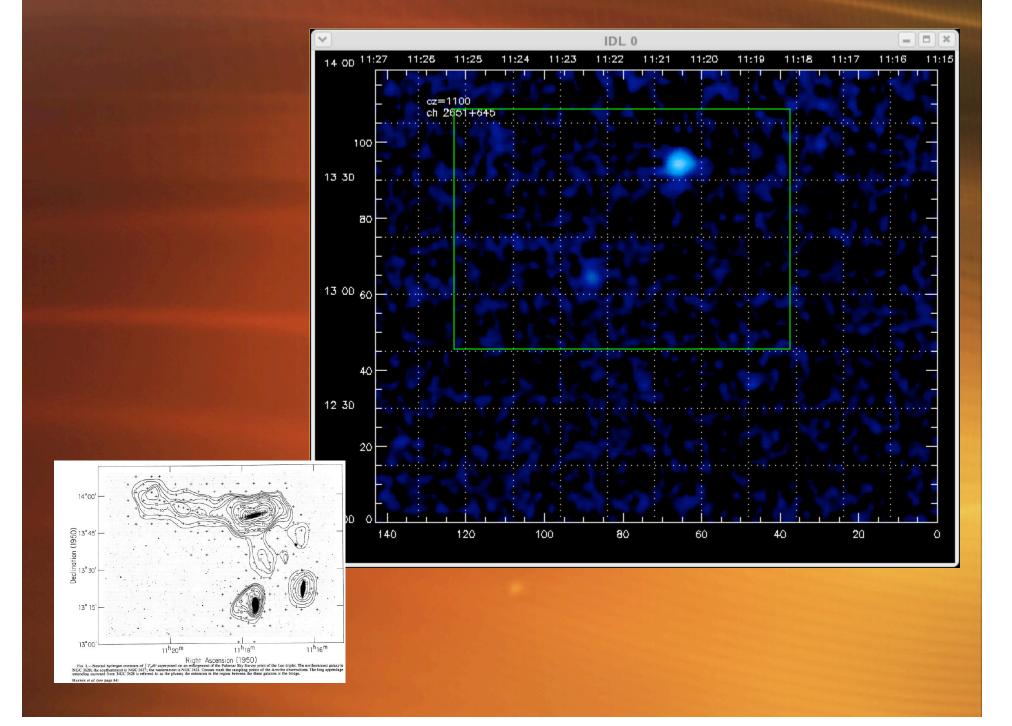
Right Ascension (1950)

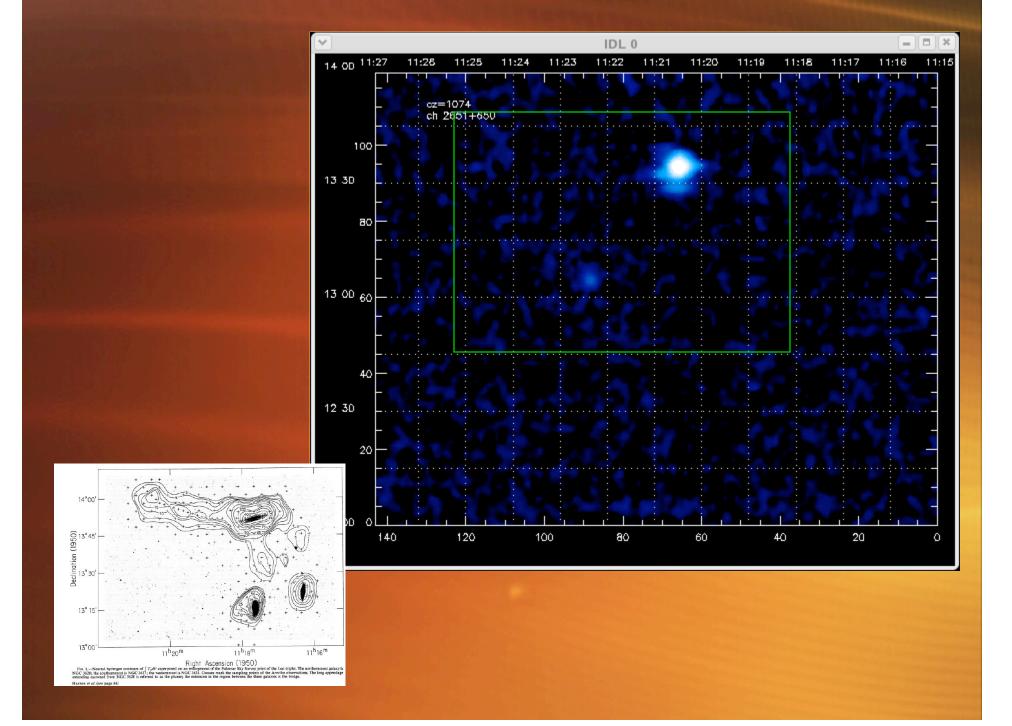
Fig. 1.—Neutral hydrogen contours of  $\int T_a dV$  superposed on an enlargement of the Palomar Sky Survey print of the Leo triplet. The northernmost galaxy is NGC 3628; the southernmost is NGC 3627; the westernmost is NGC 3623. Crosses mark the sampling points of the Arecibo observations. The long appendage extending eastward from NGC 3628 is referred to as the plume; the extension in the region between the three galaxies is the bridge.

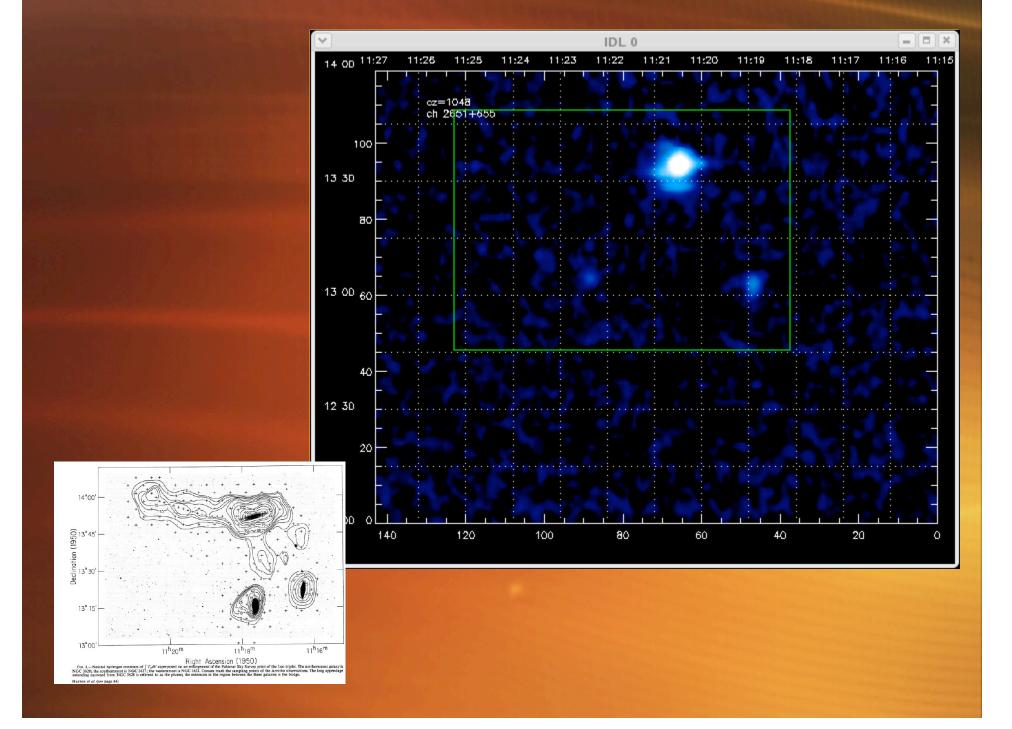
# ALFALFA Coverage of Triplet

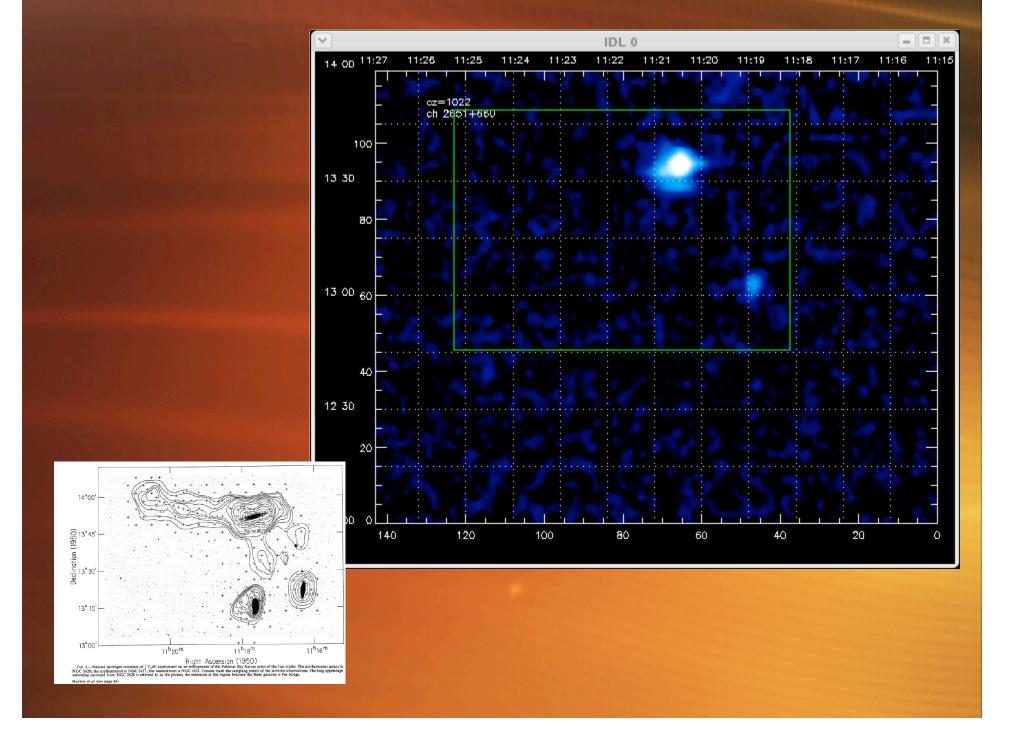


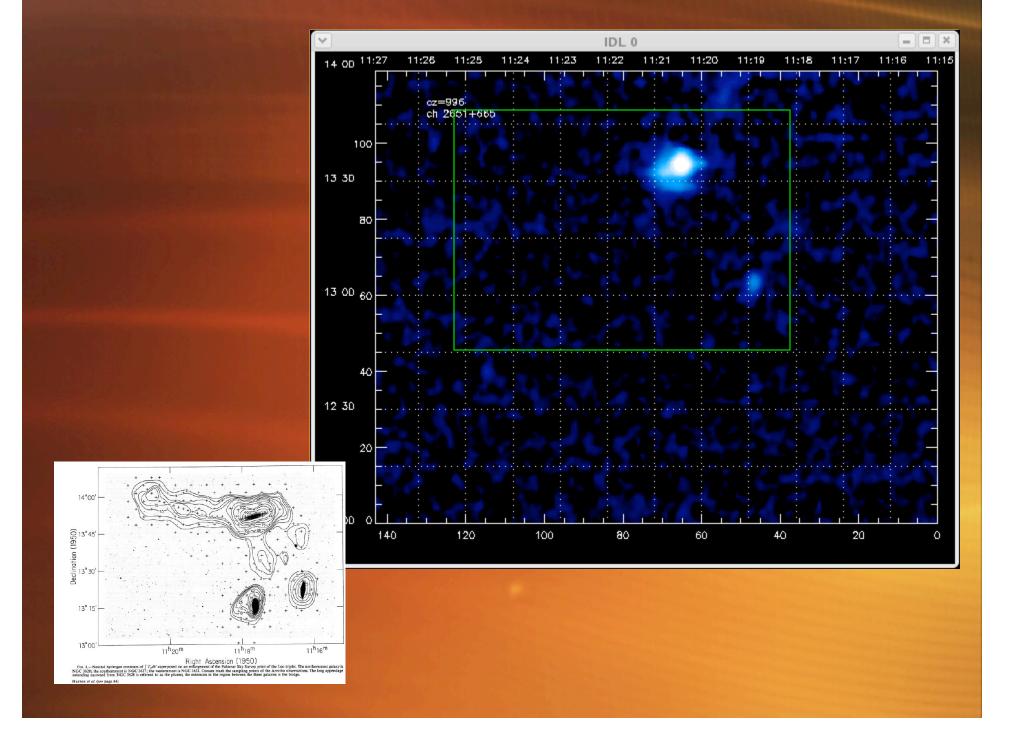


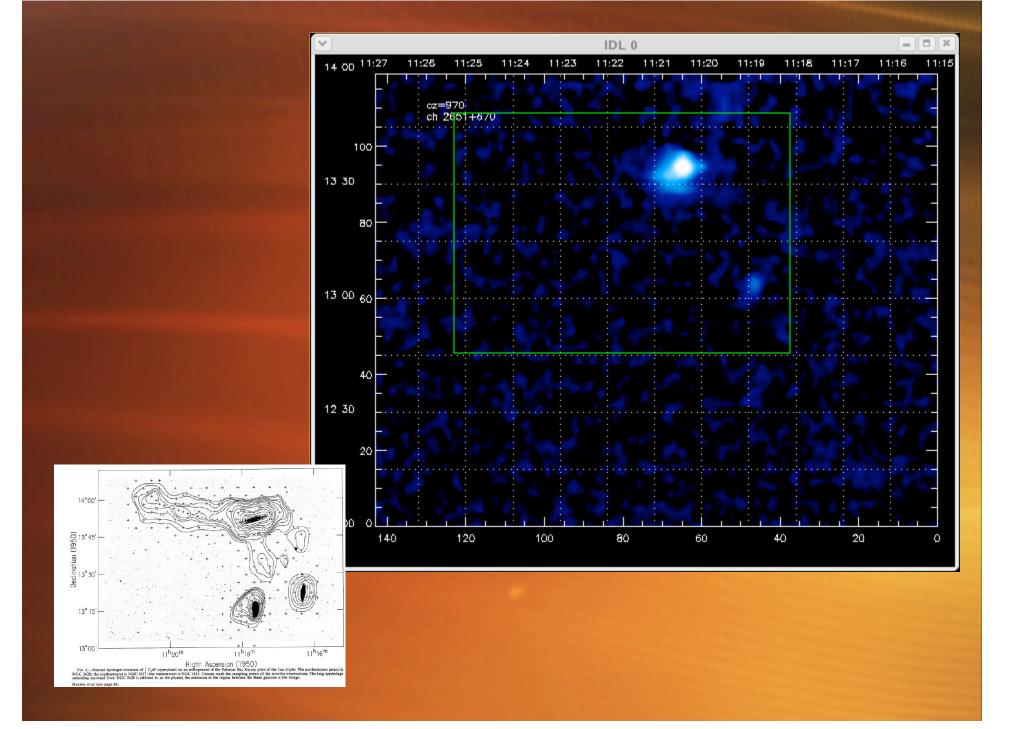


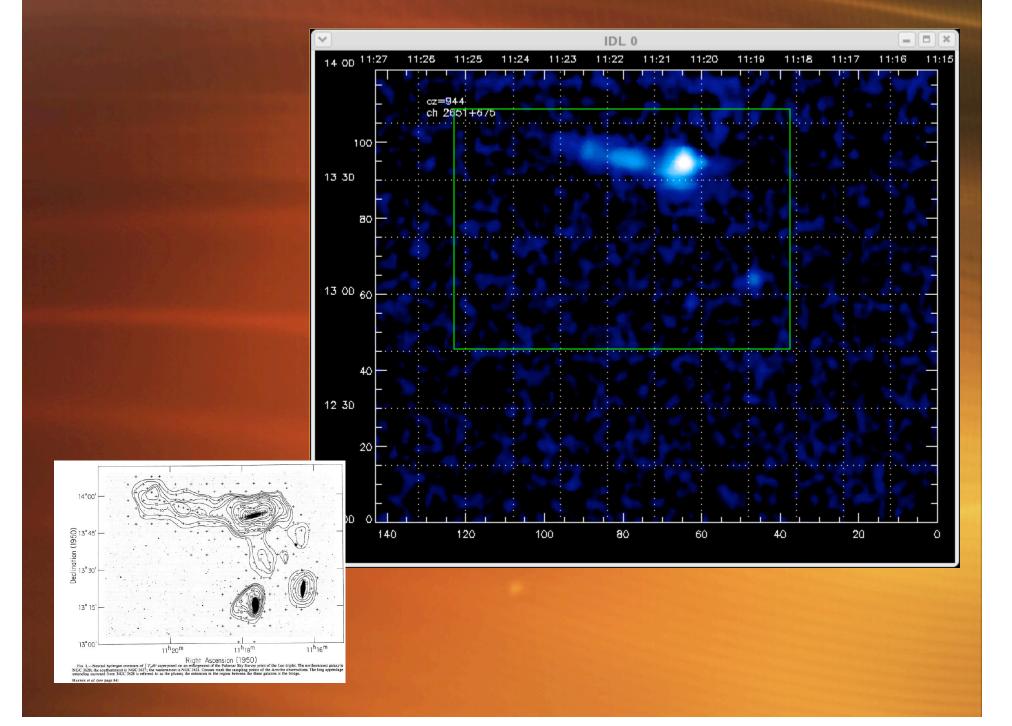


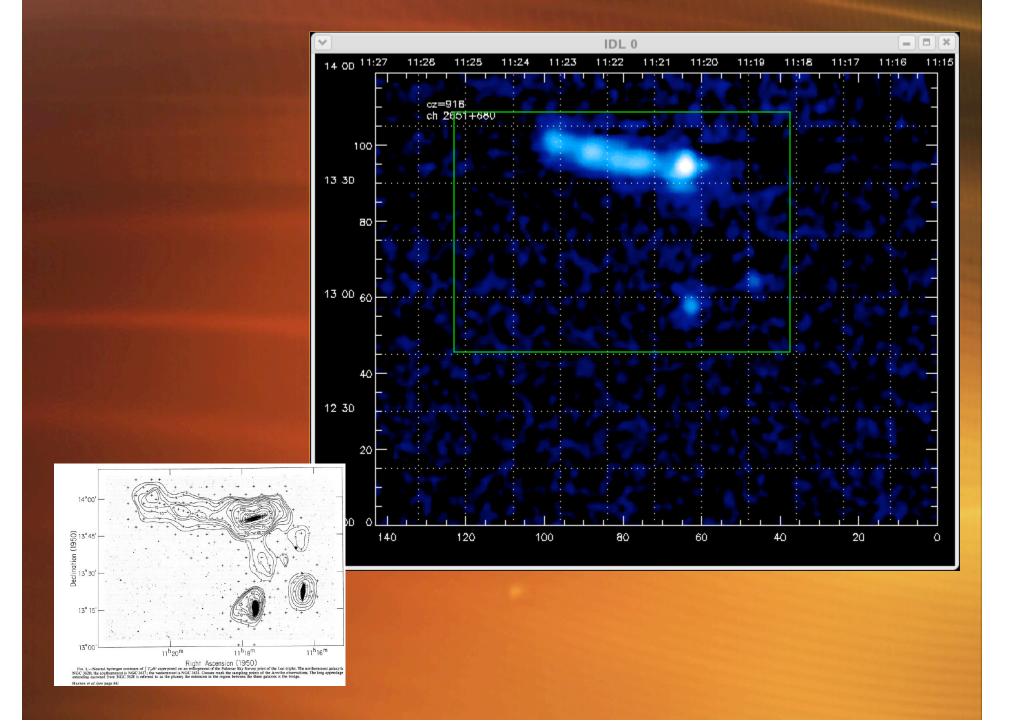


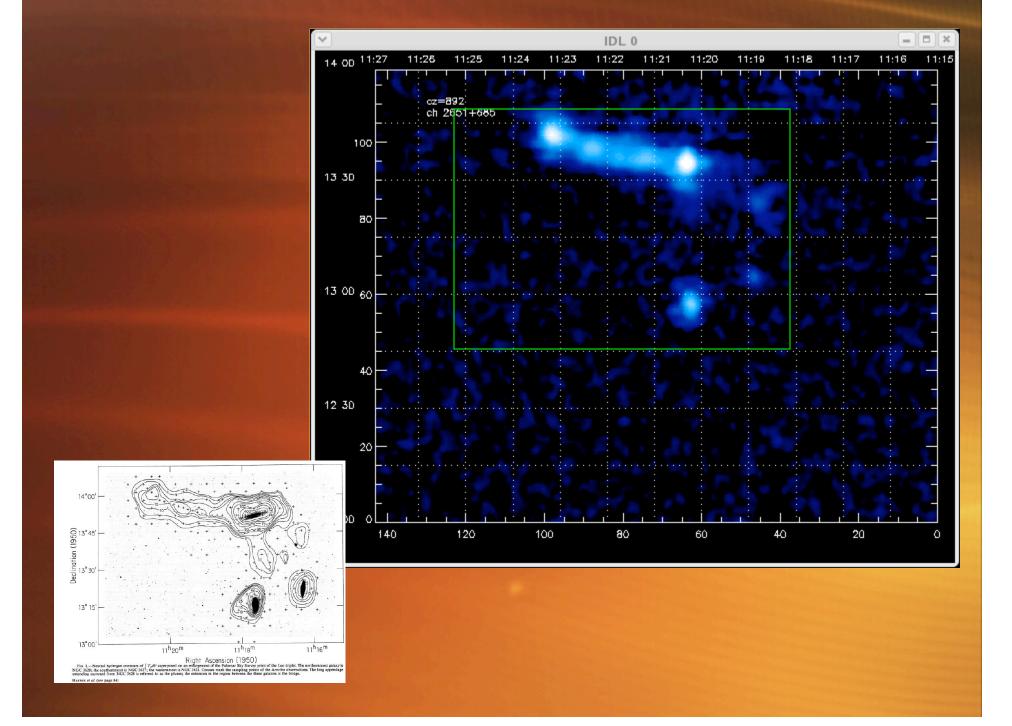


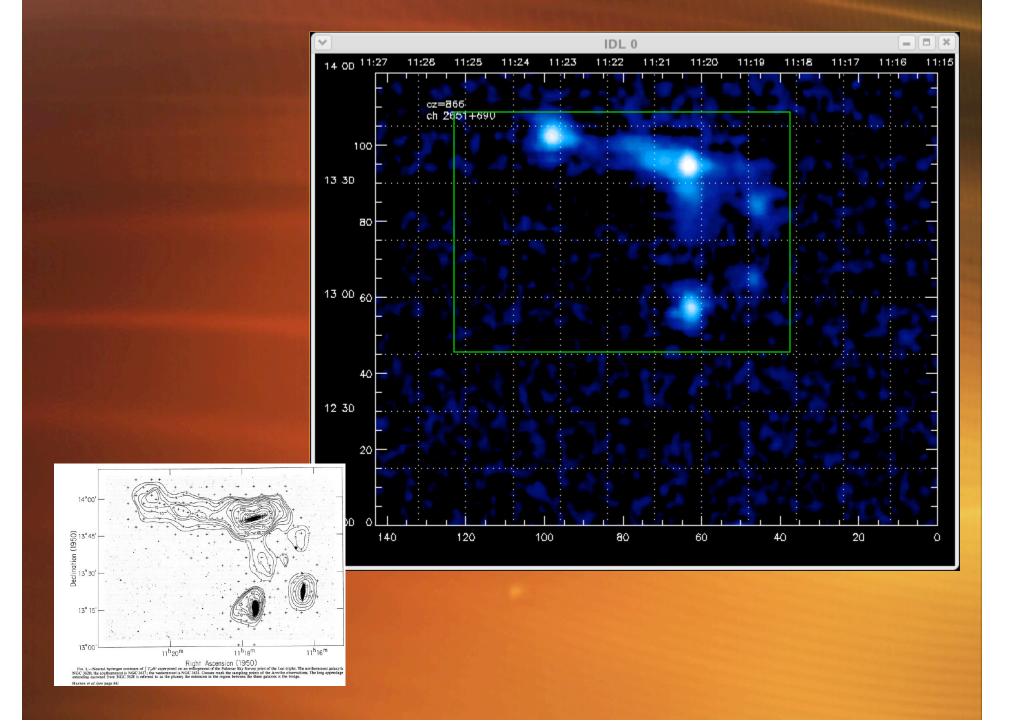


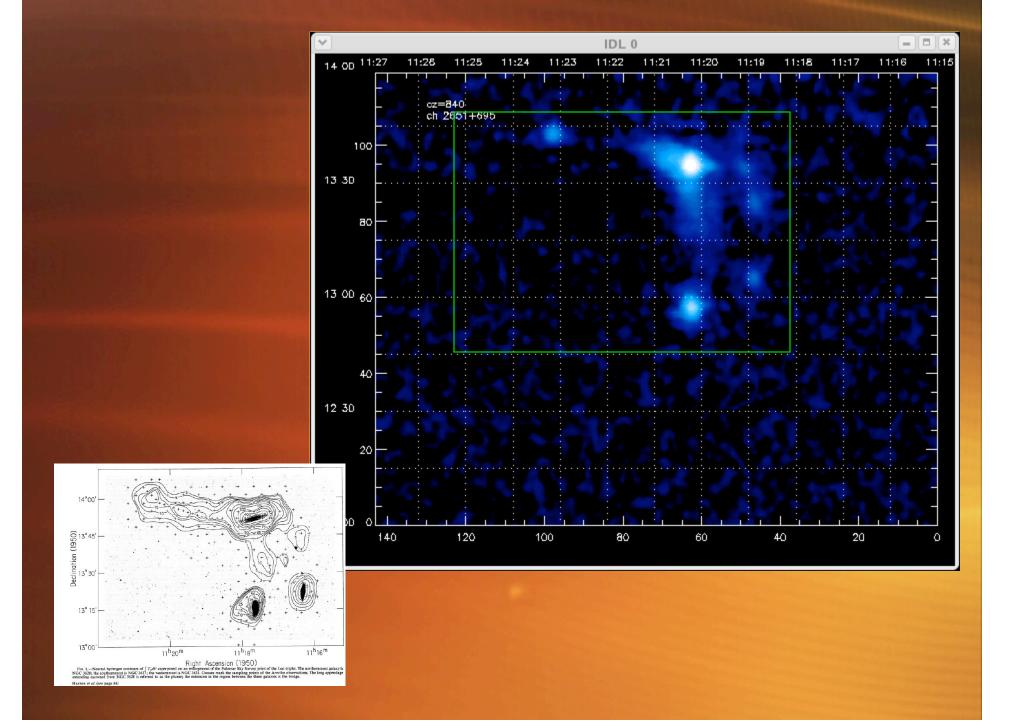


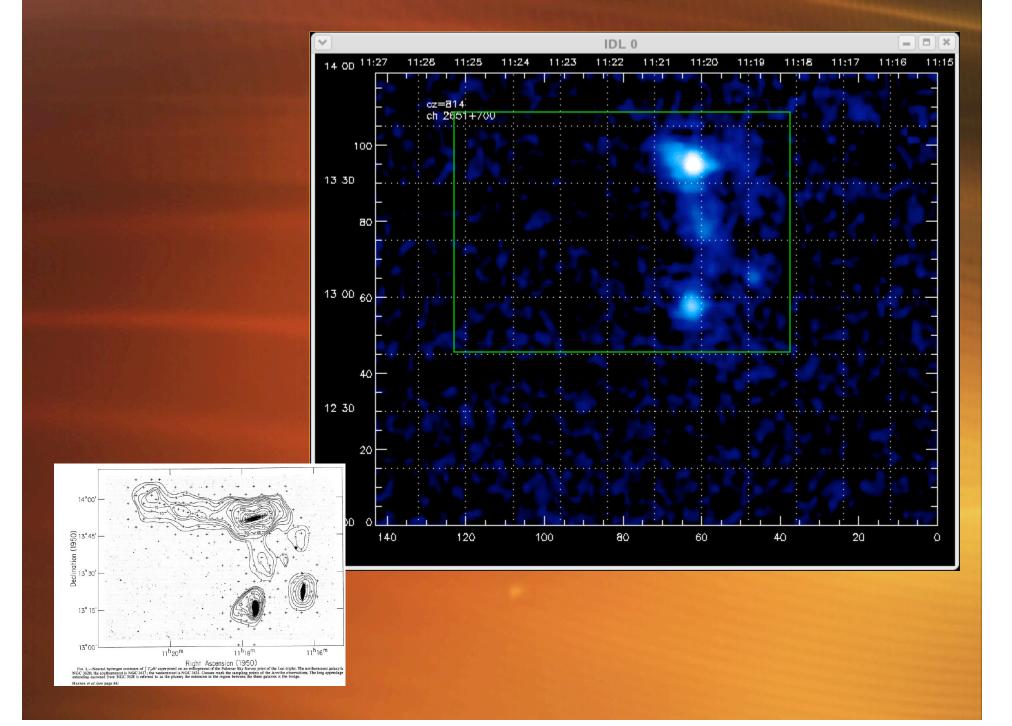


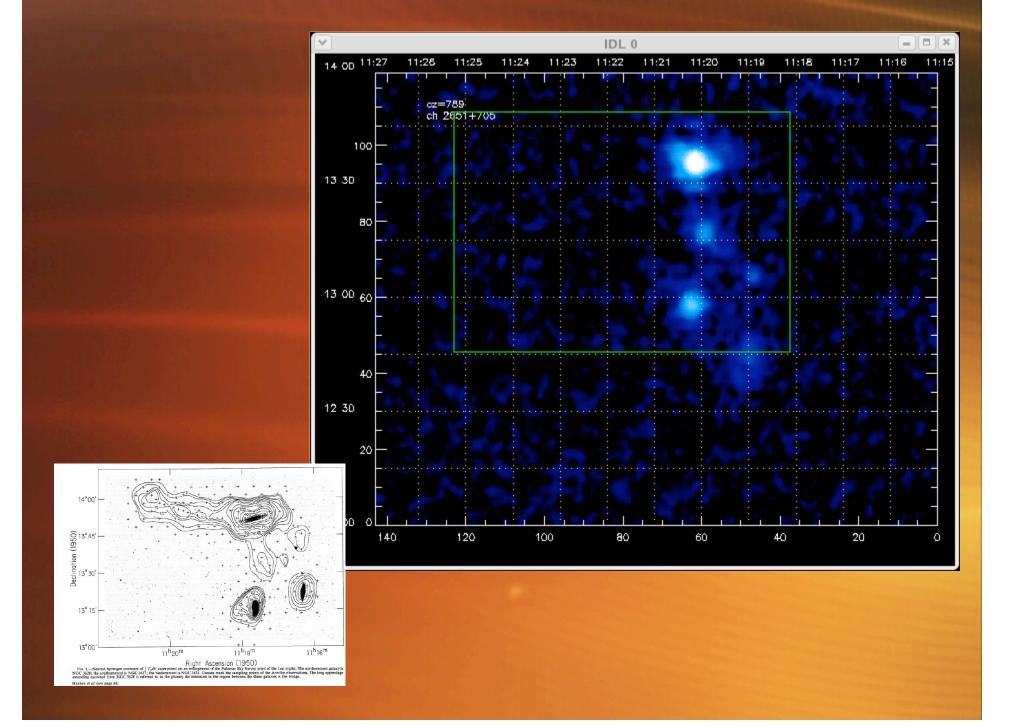


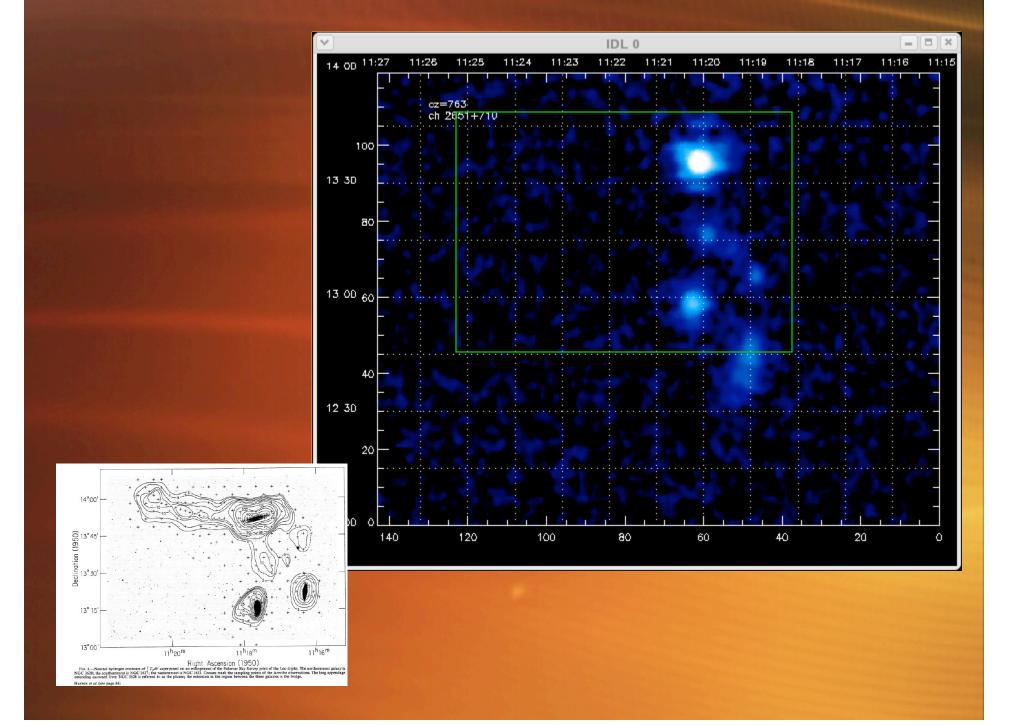


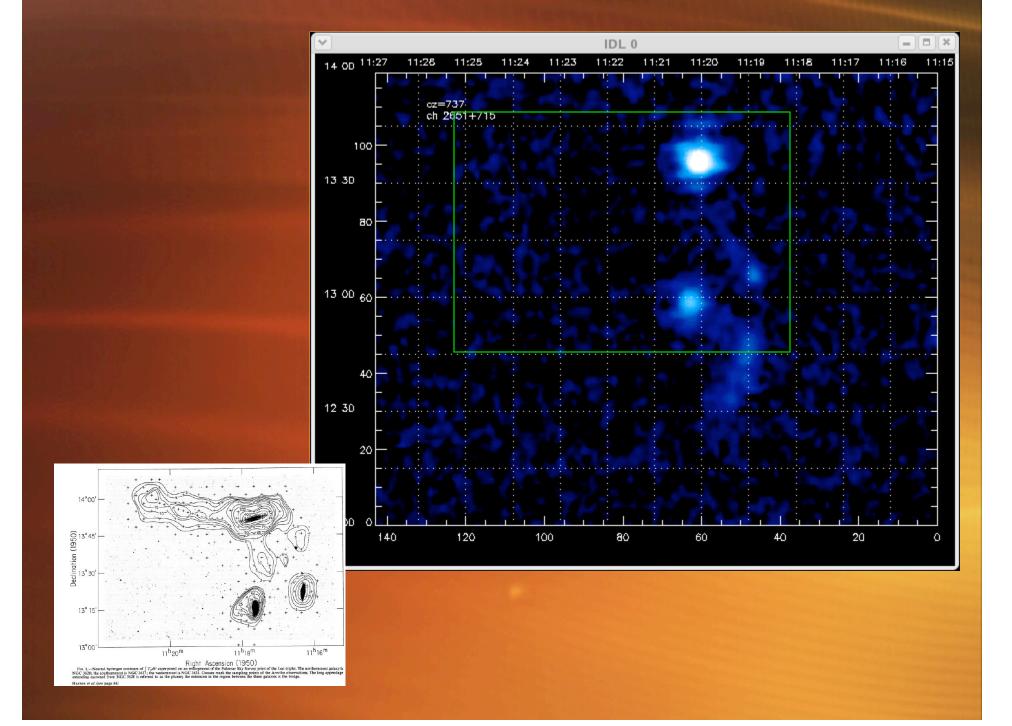


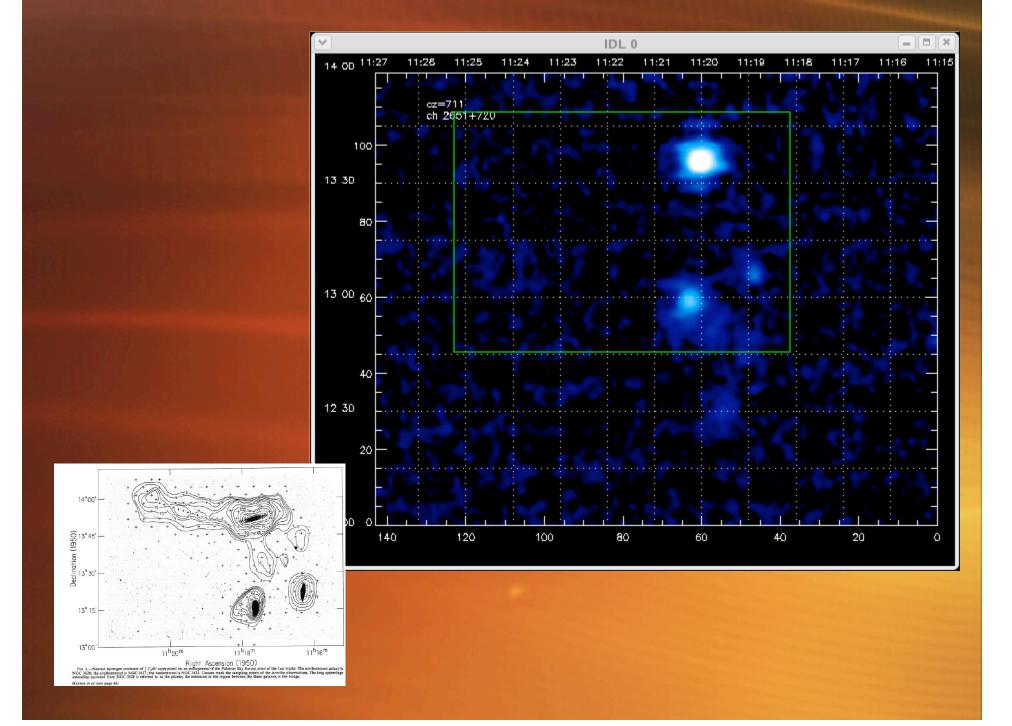


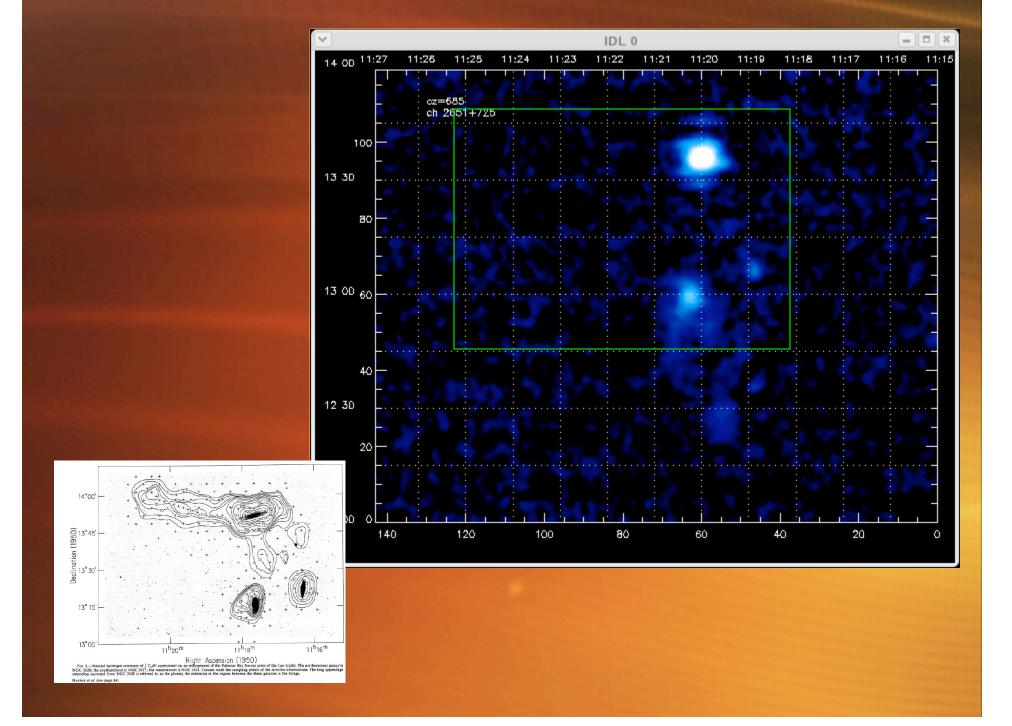


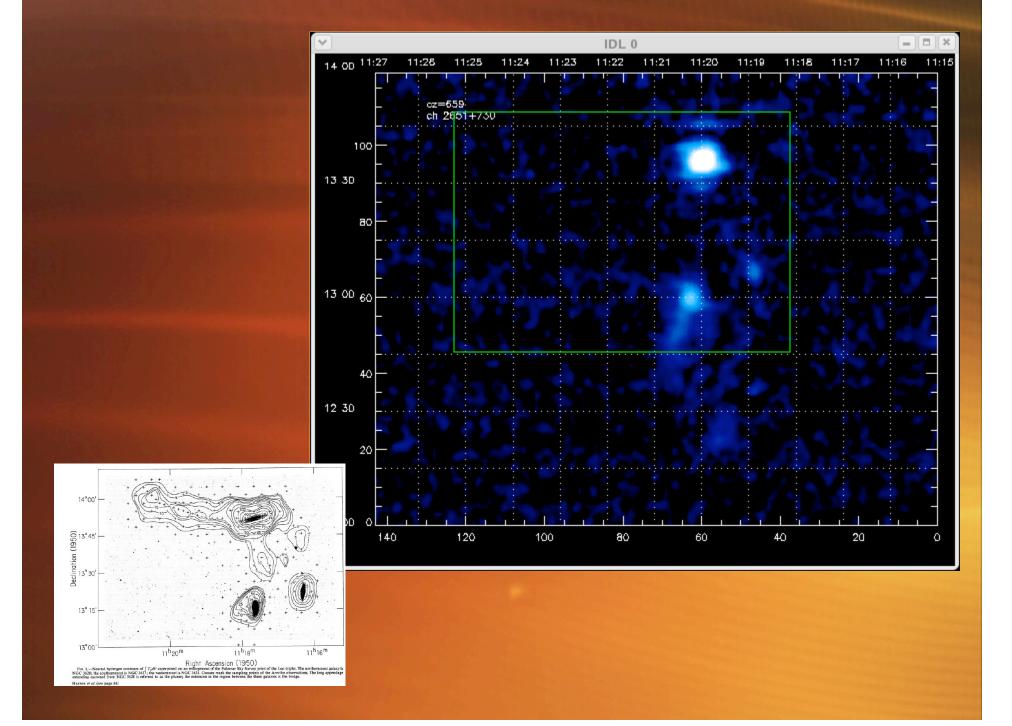


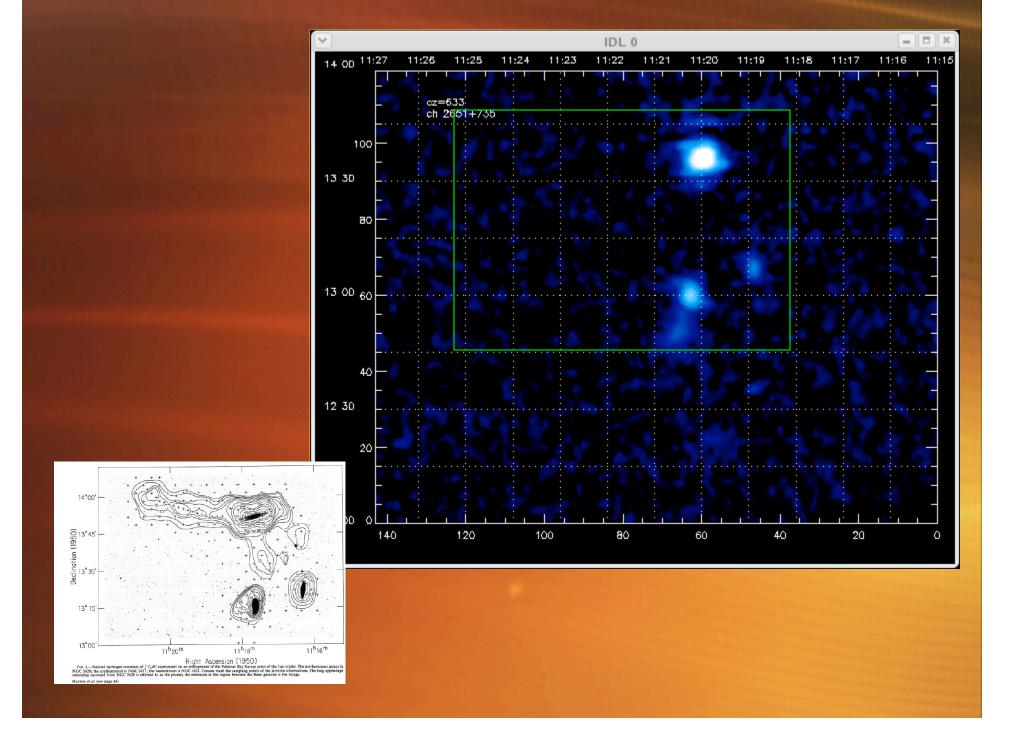


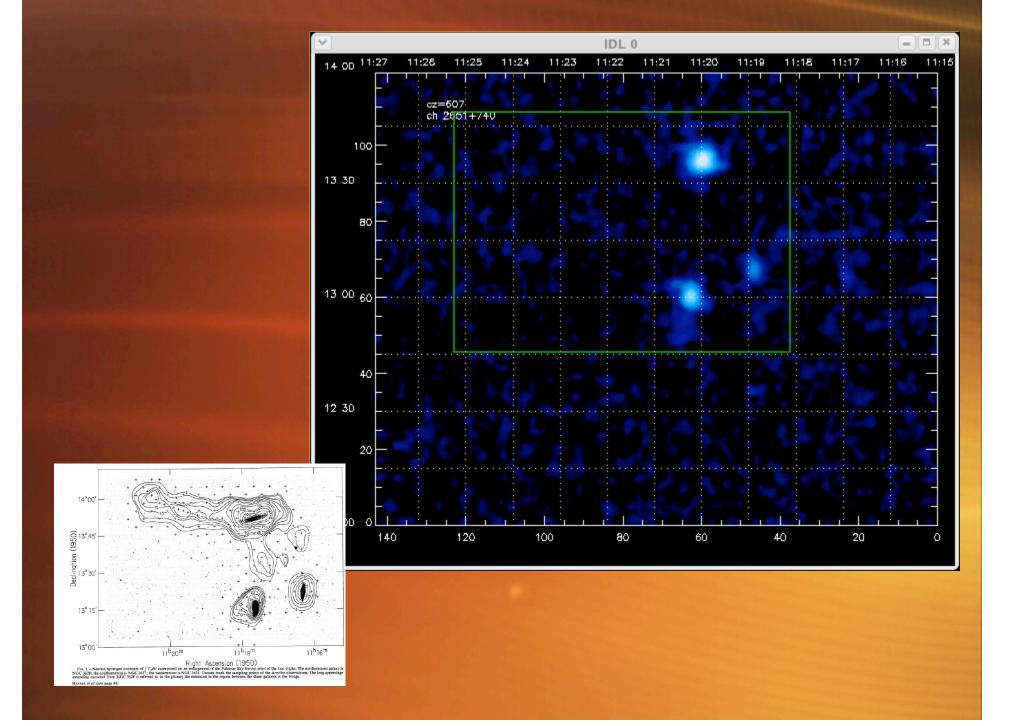


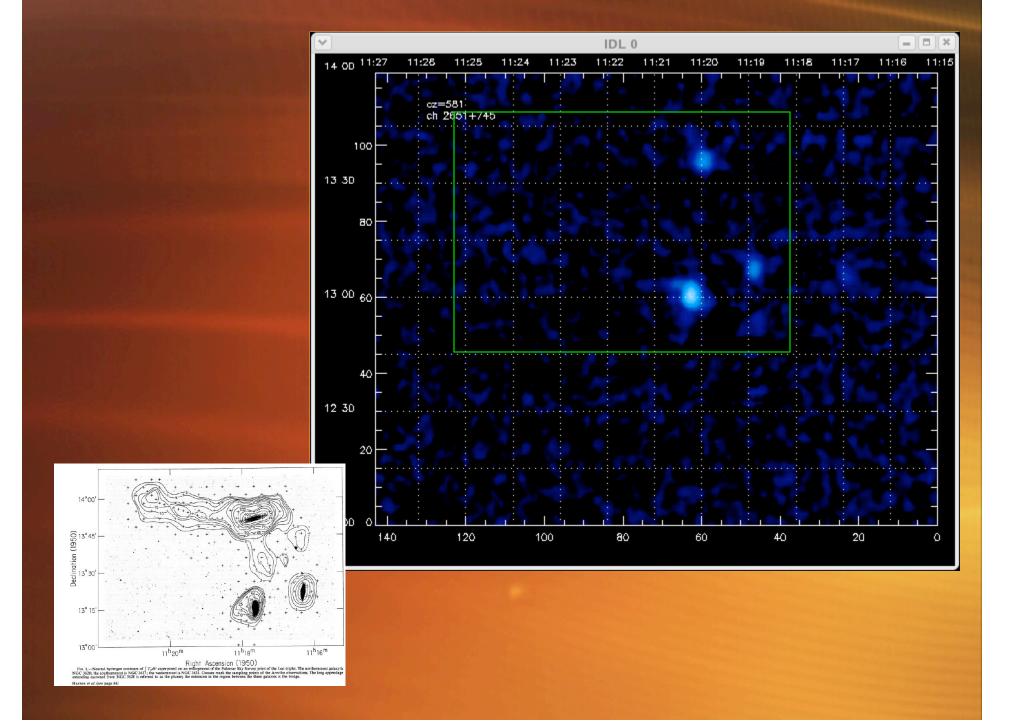


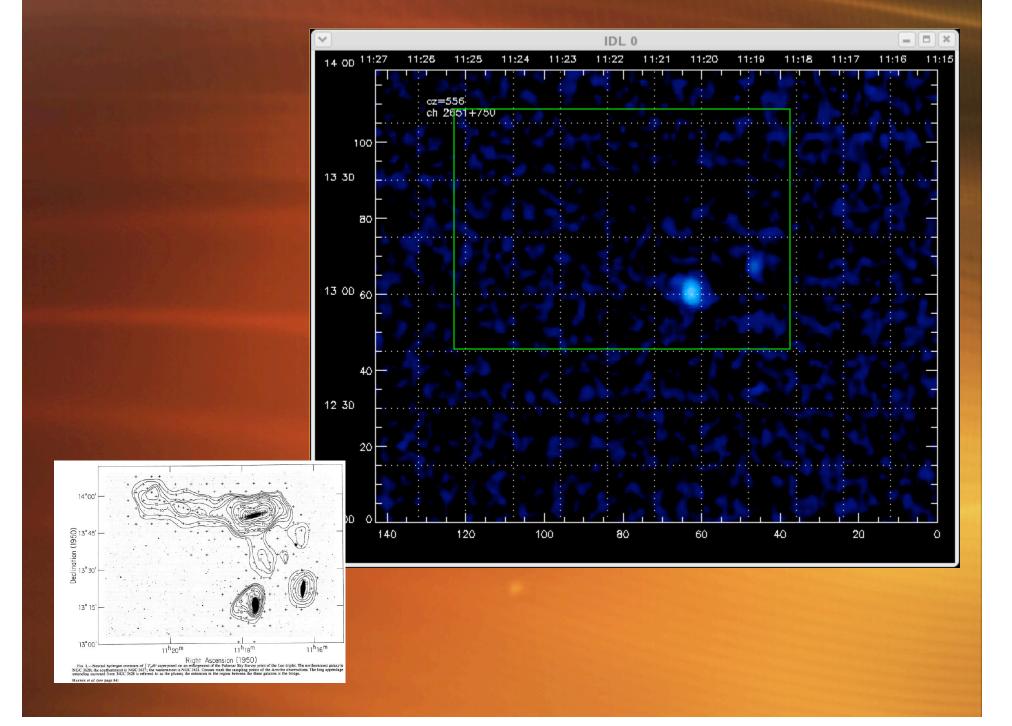


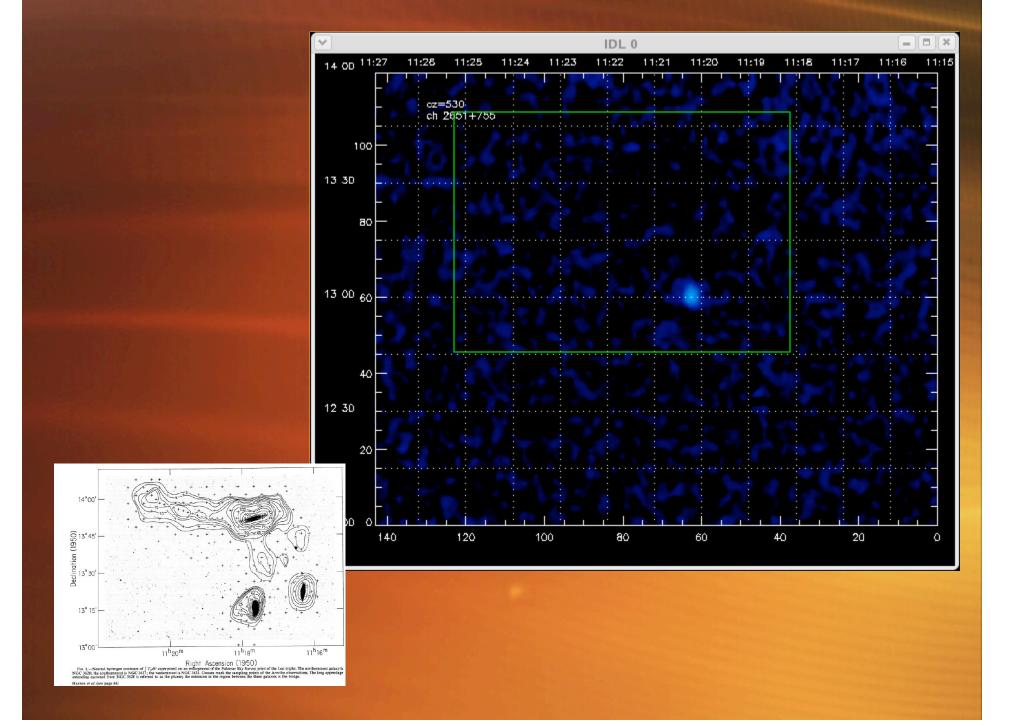


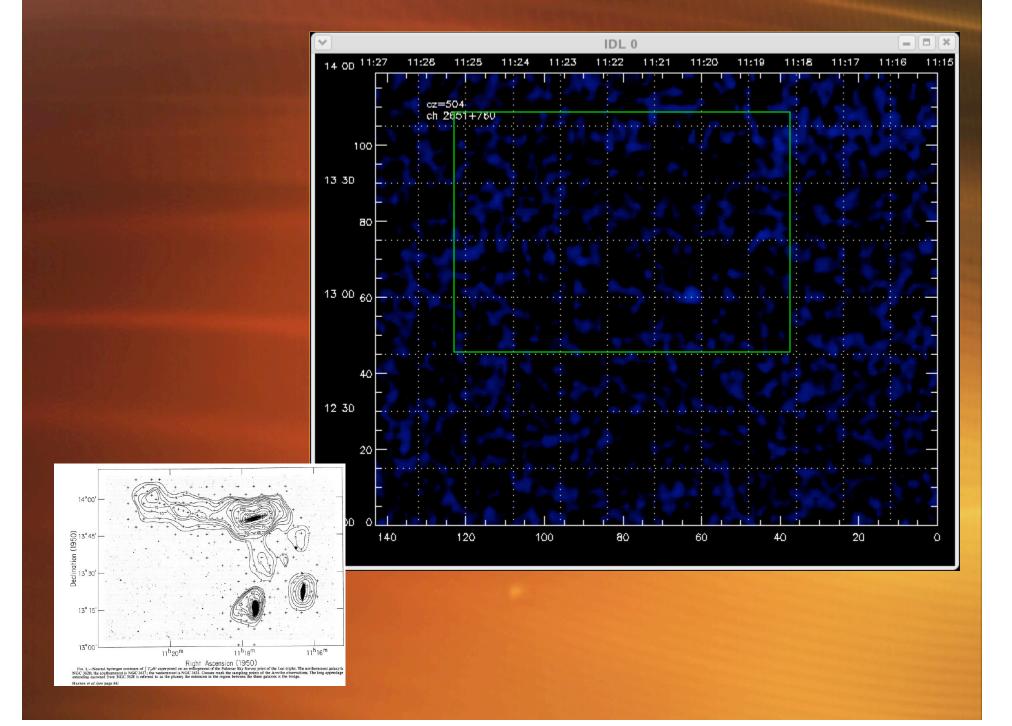




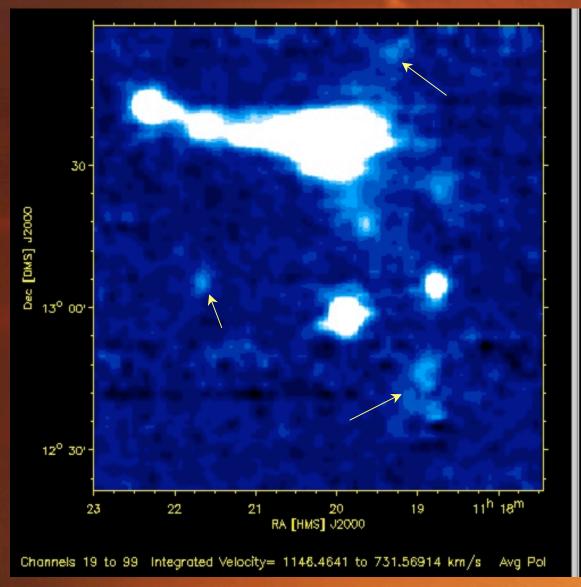


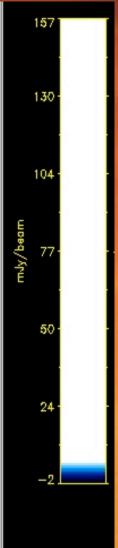






## The Leo Triplet in ALFALFA





Arrows
indicate new
HI gas clouds
not seen in
previous HI
mappings of
the region by
Haynes,
Giovanelli &
Roberts in 1979

#### HI Clouds (No Optical Counterparts)

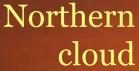
111932.6+135208

Half iso leve

Other

1400

1400





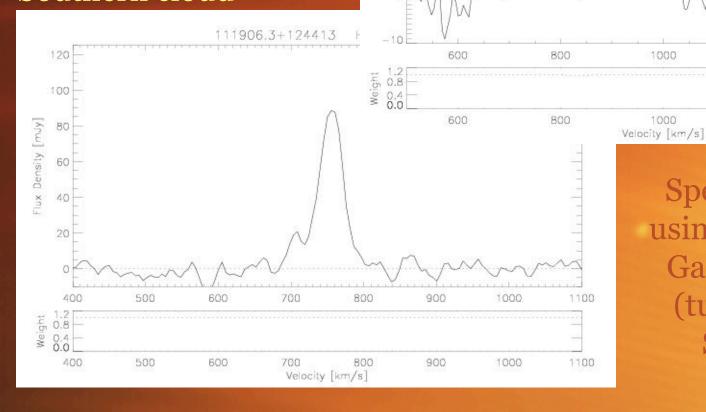
30 -

20 -

10

Flux Density [mdy]

#### Southern cloud



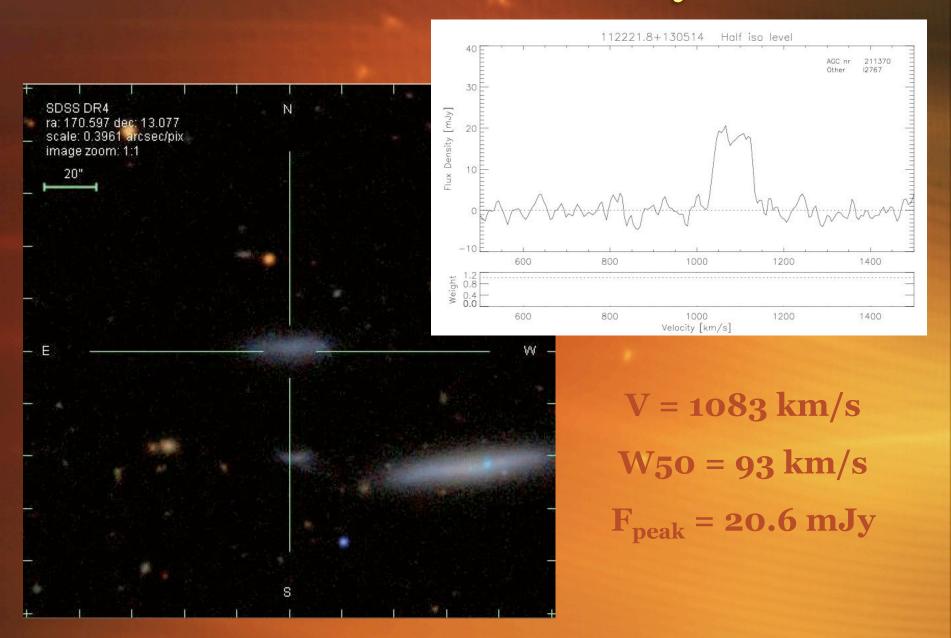
Spectra extracted using LOVEDATA's GalFlux Program (tune back in on Saturday ...)

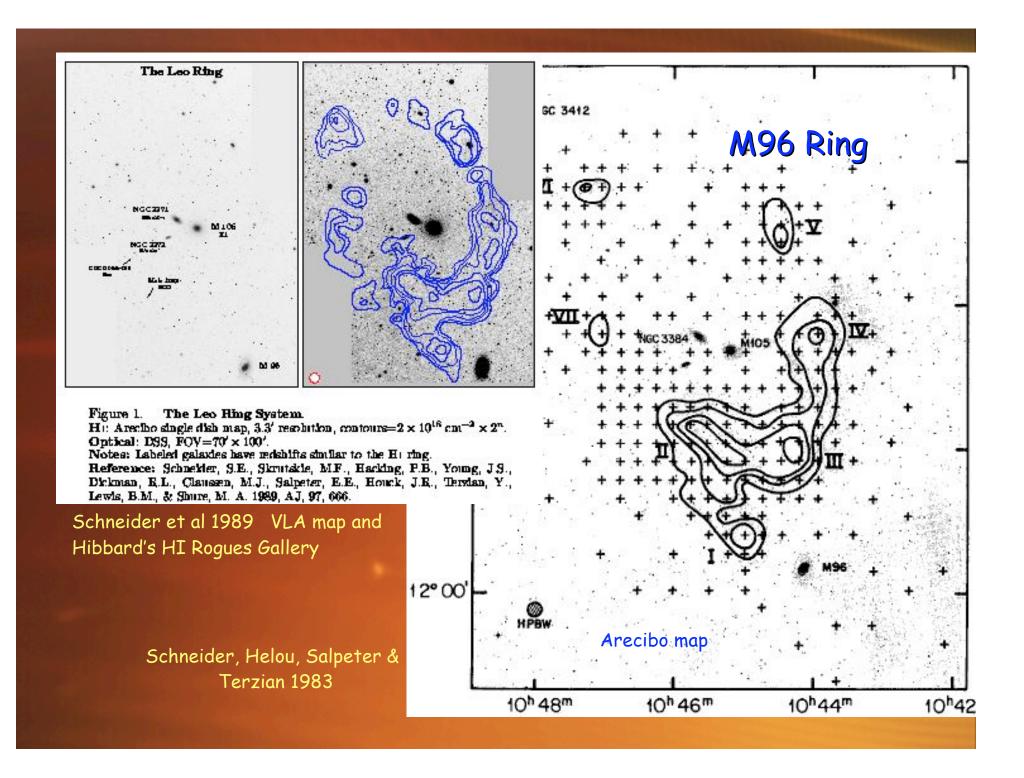
1200

1200

1000

## New Dwarf Galaxy!



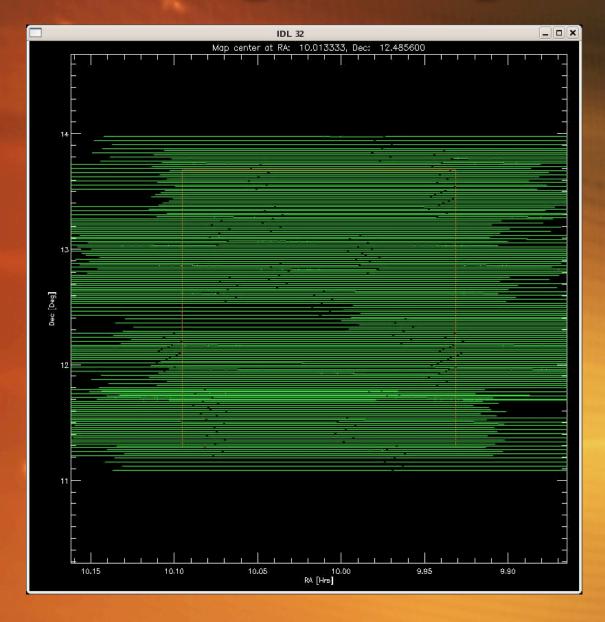


# **ALFALFA Coverage of Ring**

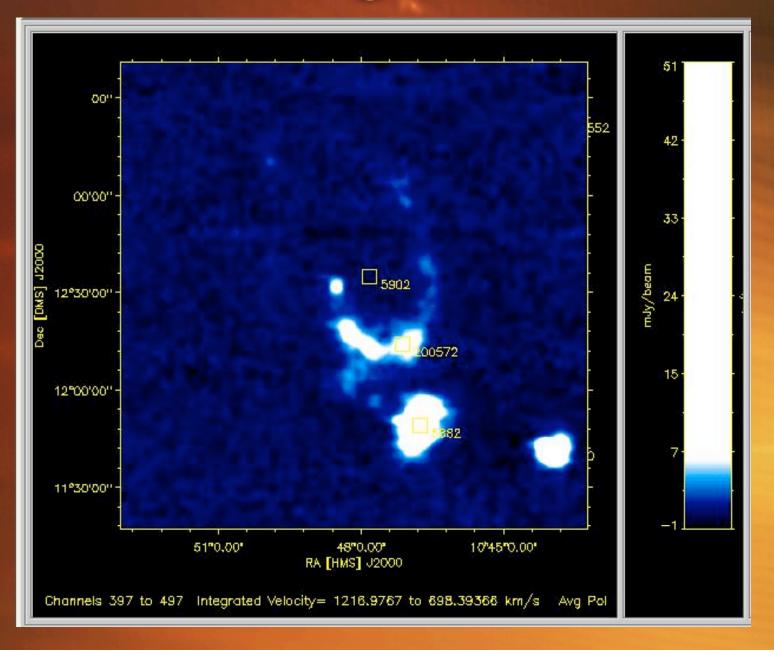
Grid Center:

RA 10<sup>h</sup>48<sup>m</sup>08<sup>s</sup>

Dec 12<sup>o</sup>29'08"



## The Leo Ring in ALFALFA



#### What's Next?

- Continue making source catalogs for the Leo region (that's 60 ALFALFA grids - not an easy job!)
- Determine group status and statistics among the several galaxy groups in the Leo region
- Follow up on objects with no obvious optical counterparts (what are they?)
- Determine the HIMF for this unique environment
- And as always ... take more data!



# Optically-Selected Sample

